

Achromatic multi-four quadrant phase mask : laboratory demonstration

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Lyot conference

October 2010



Outline

Monochromatic FQPM

Multi-FQPM principle

Performance in full pupil

Performance with obscuration

Conclusions

Outline

Monochromatic FQPM

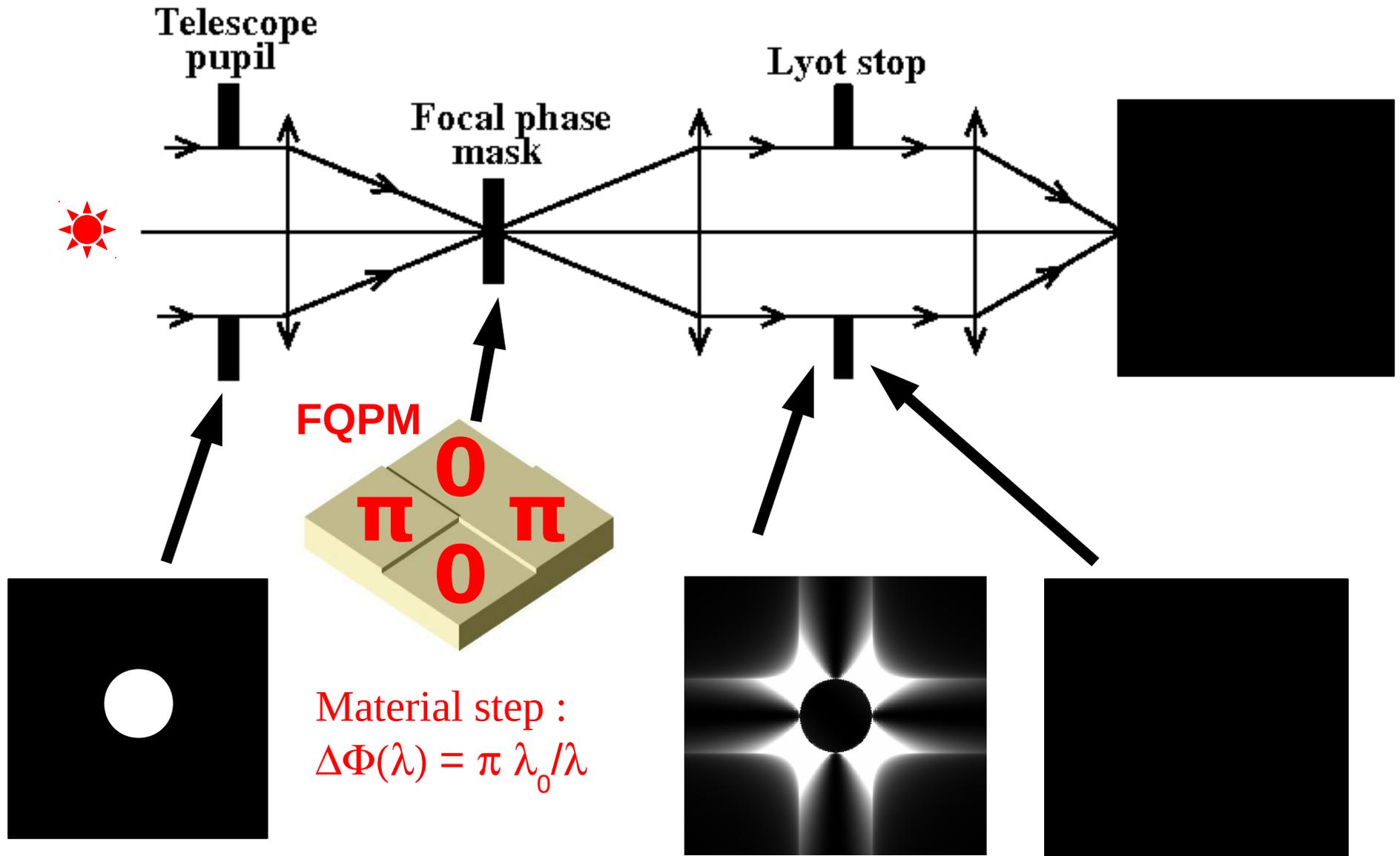
Multi-FQPM principle

Performance in full pupil

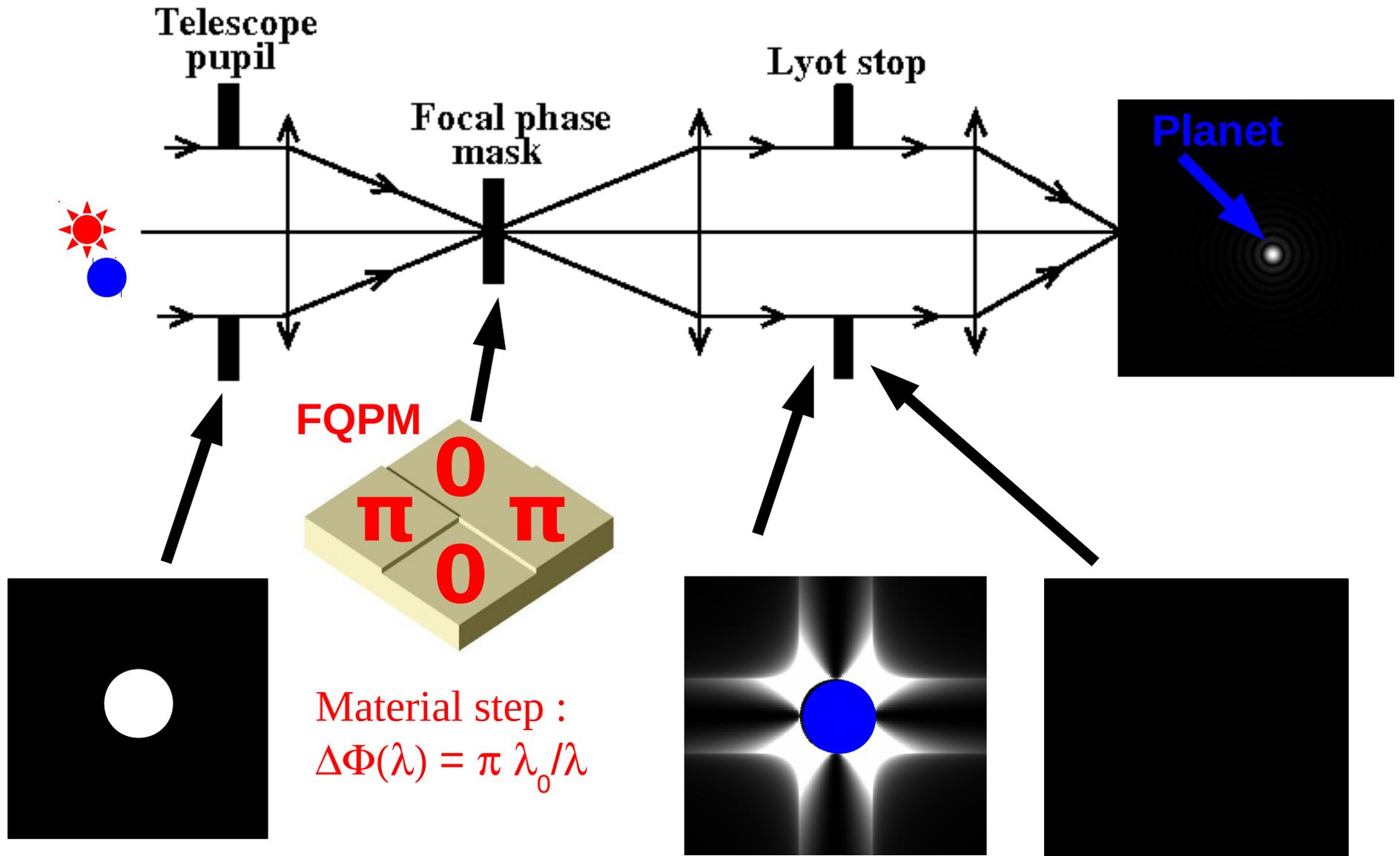
Performance with obscuration

Conclusions

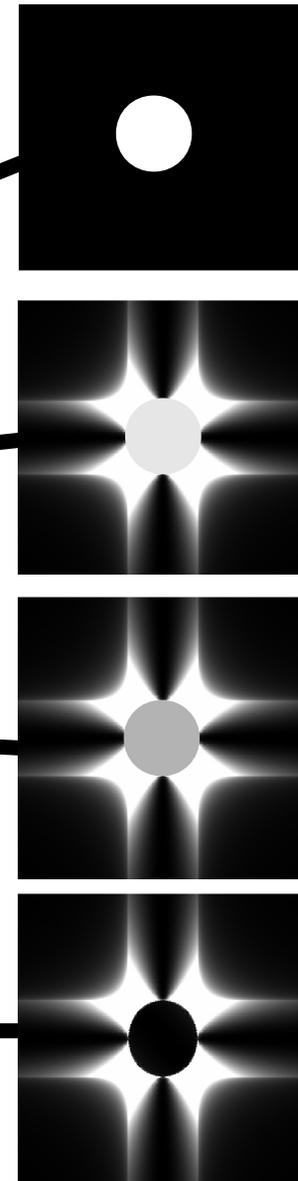
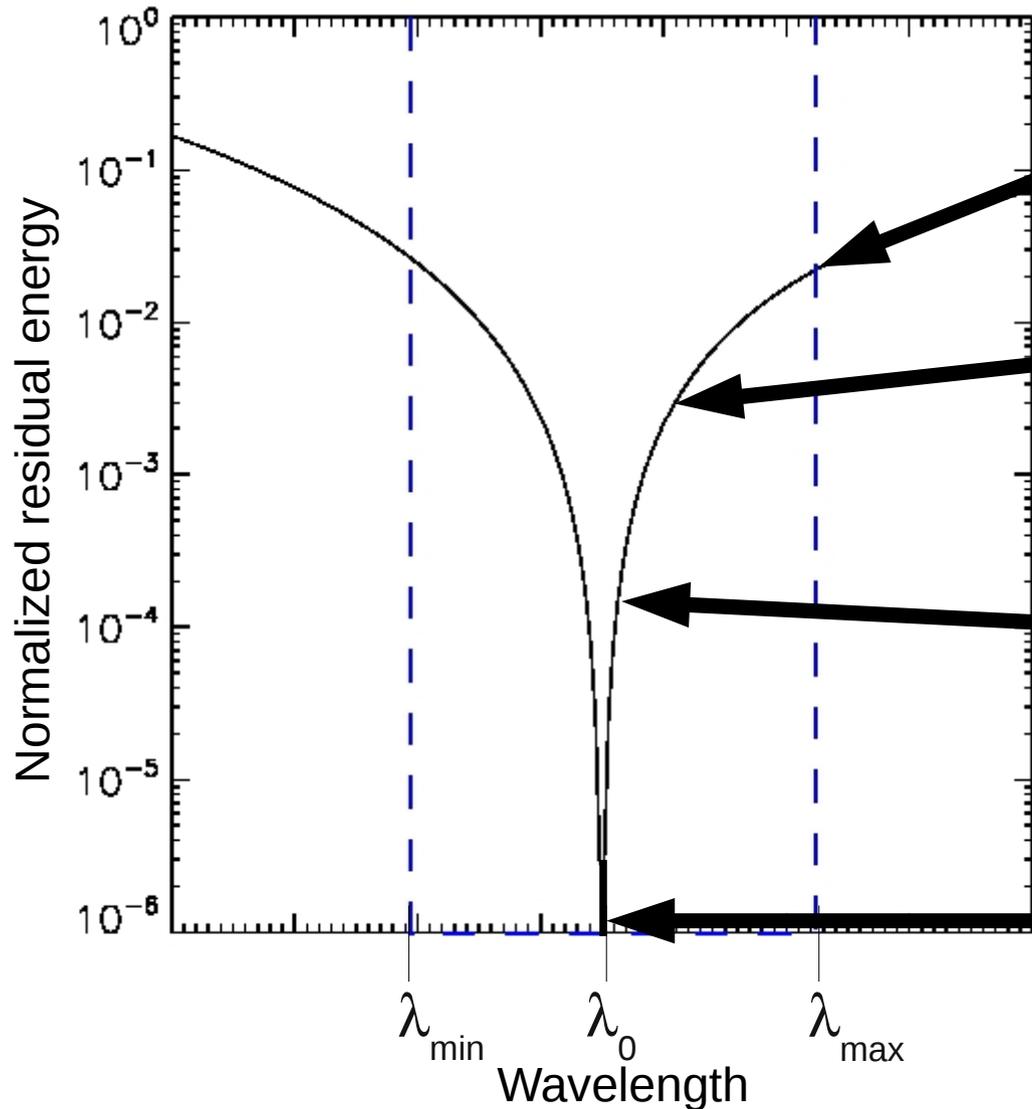
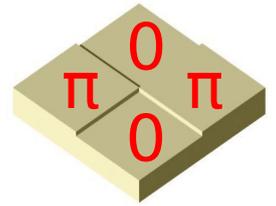
Monochromatic four quadrant phase mask



Monochromatic four quadrant phase mask



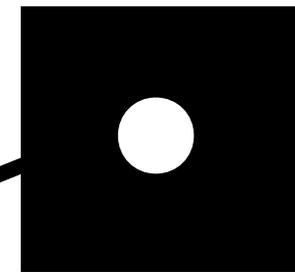
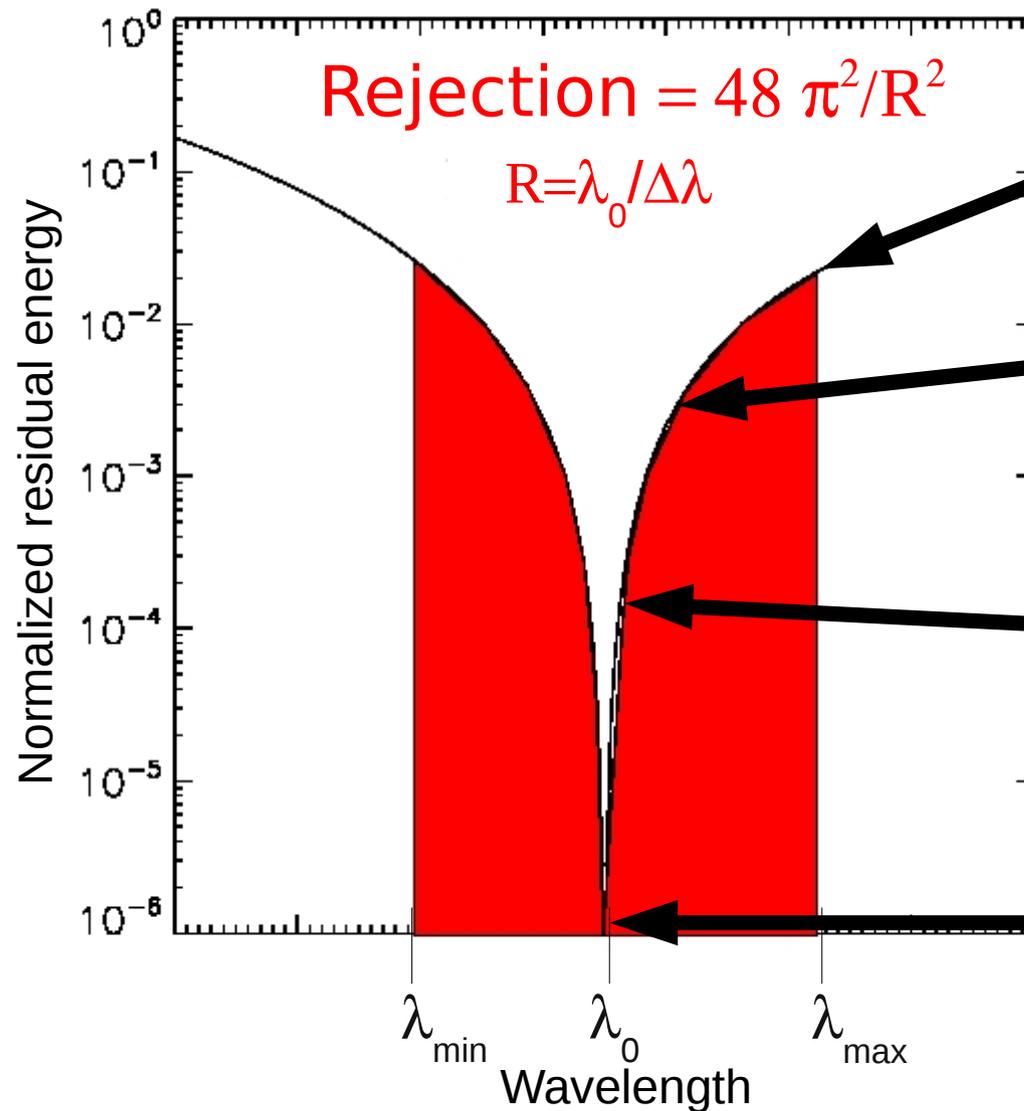
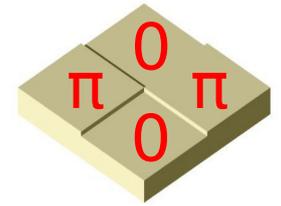
Single monochromatic FQPM



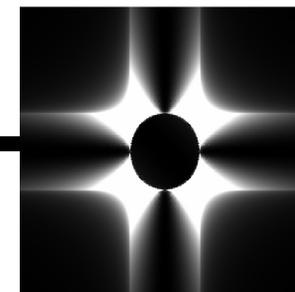
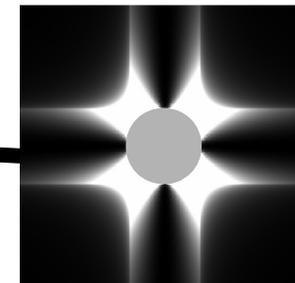
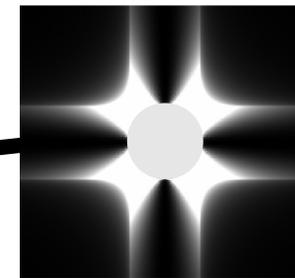
No effect

Perfect

Single monochromatic FQPM

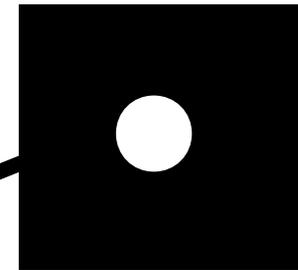
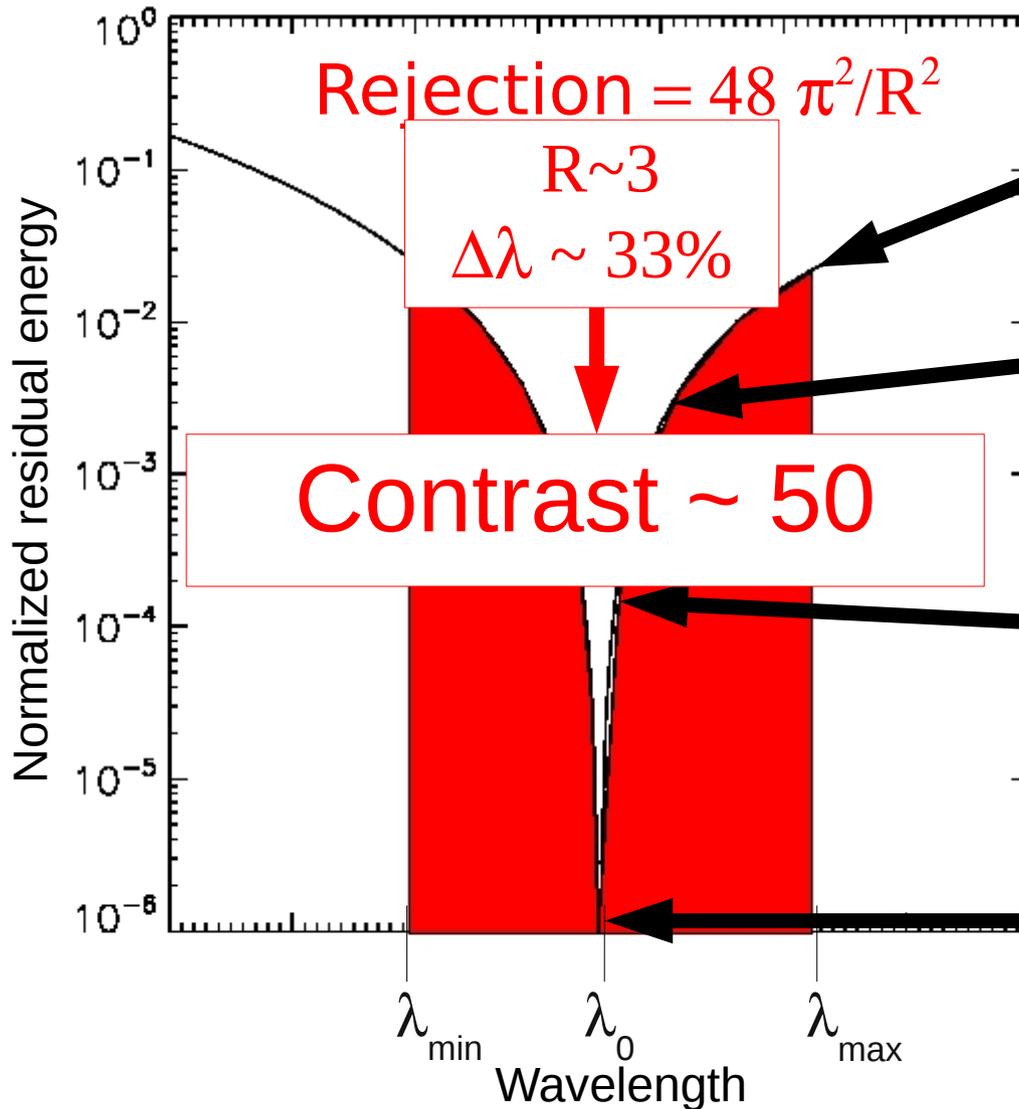
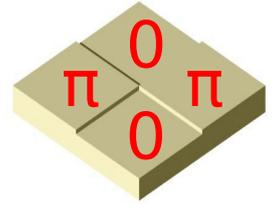


No effect

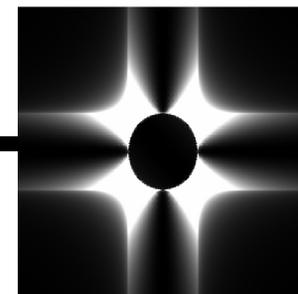
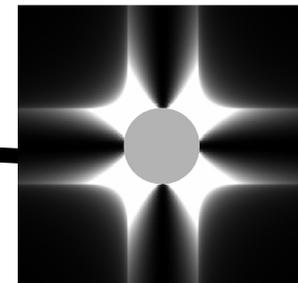
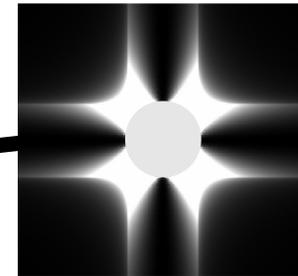


Perfect

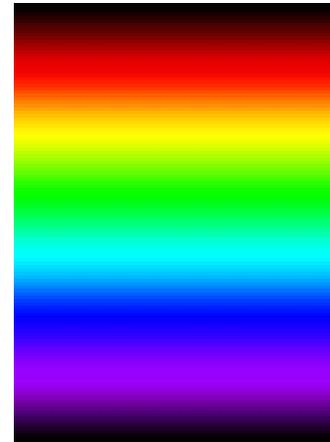
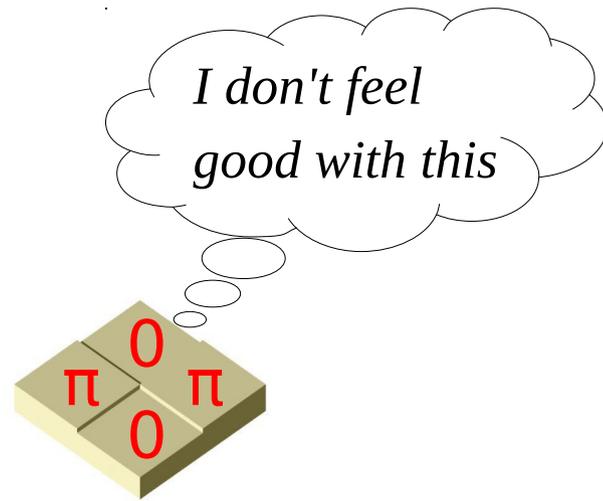
Single monochromatic FQPM



No effect



Perfect

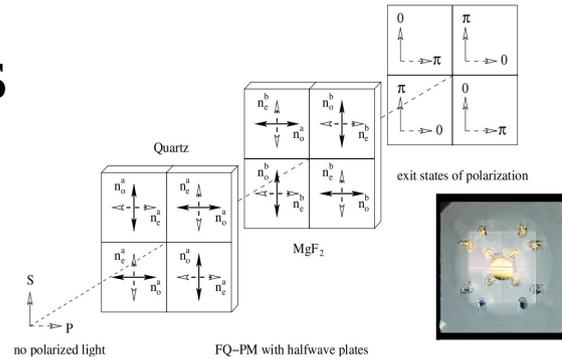


FQPM achromatization

- Halfwave plates

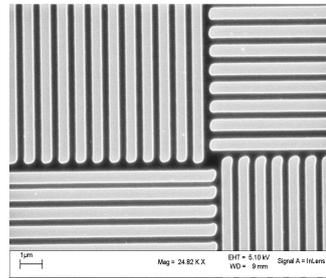
Mawet et al. 2006

Boccaletti et al. 2008



- Sub-lambda gratings

Mawet et al. 2005, 2006



- Mach-Zender interferometer

Carlotti et al. 2008

- Multi-stage four quadrant phase mask → MFQPM

Baudoz et al. 2008

Outline

Monochromatic FQPM

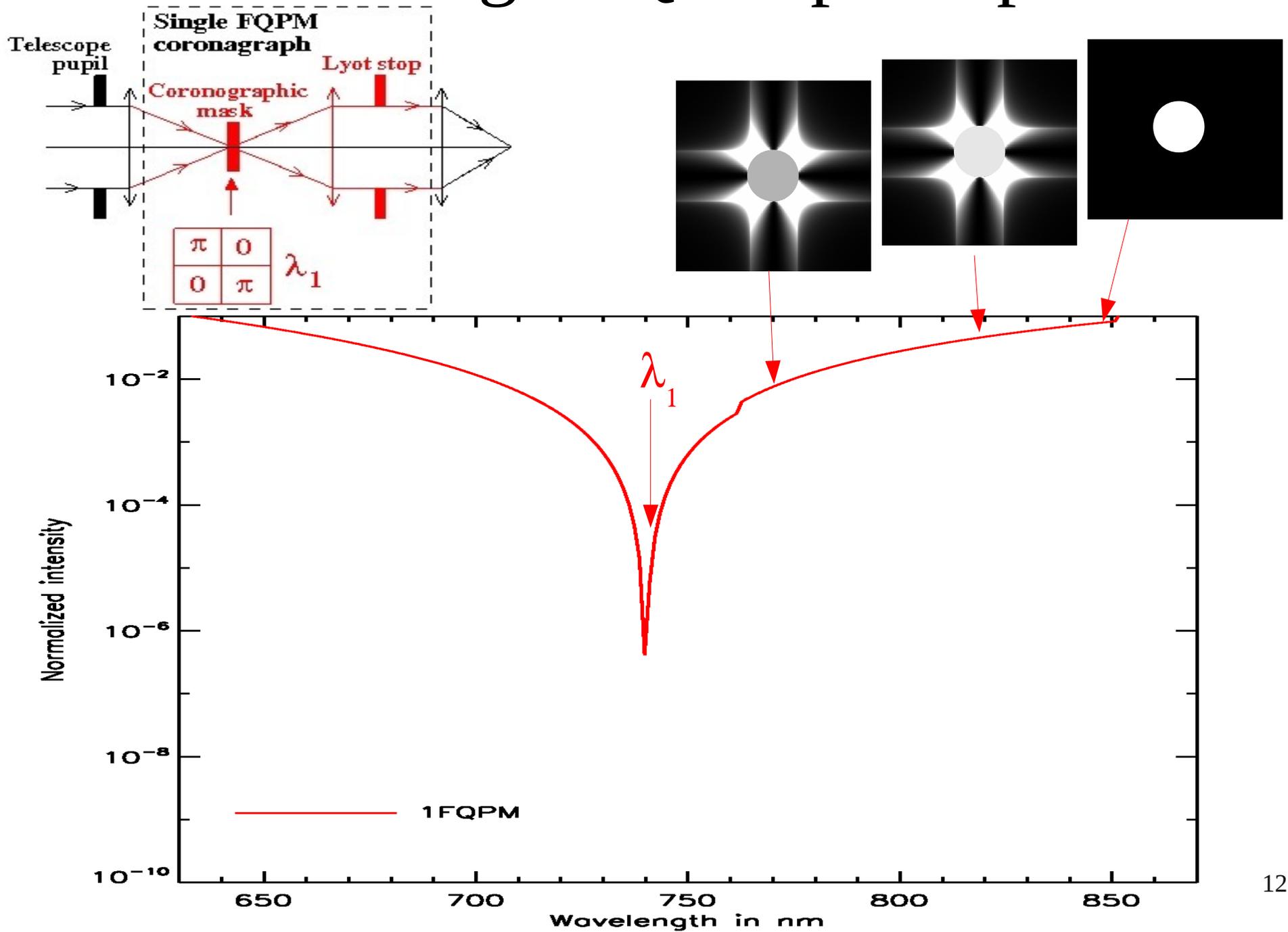
Multi-FQPM principle

Performance in full pupil

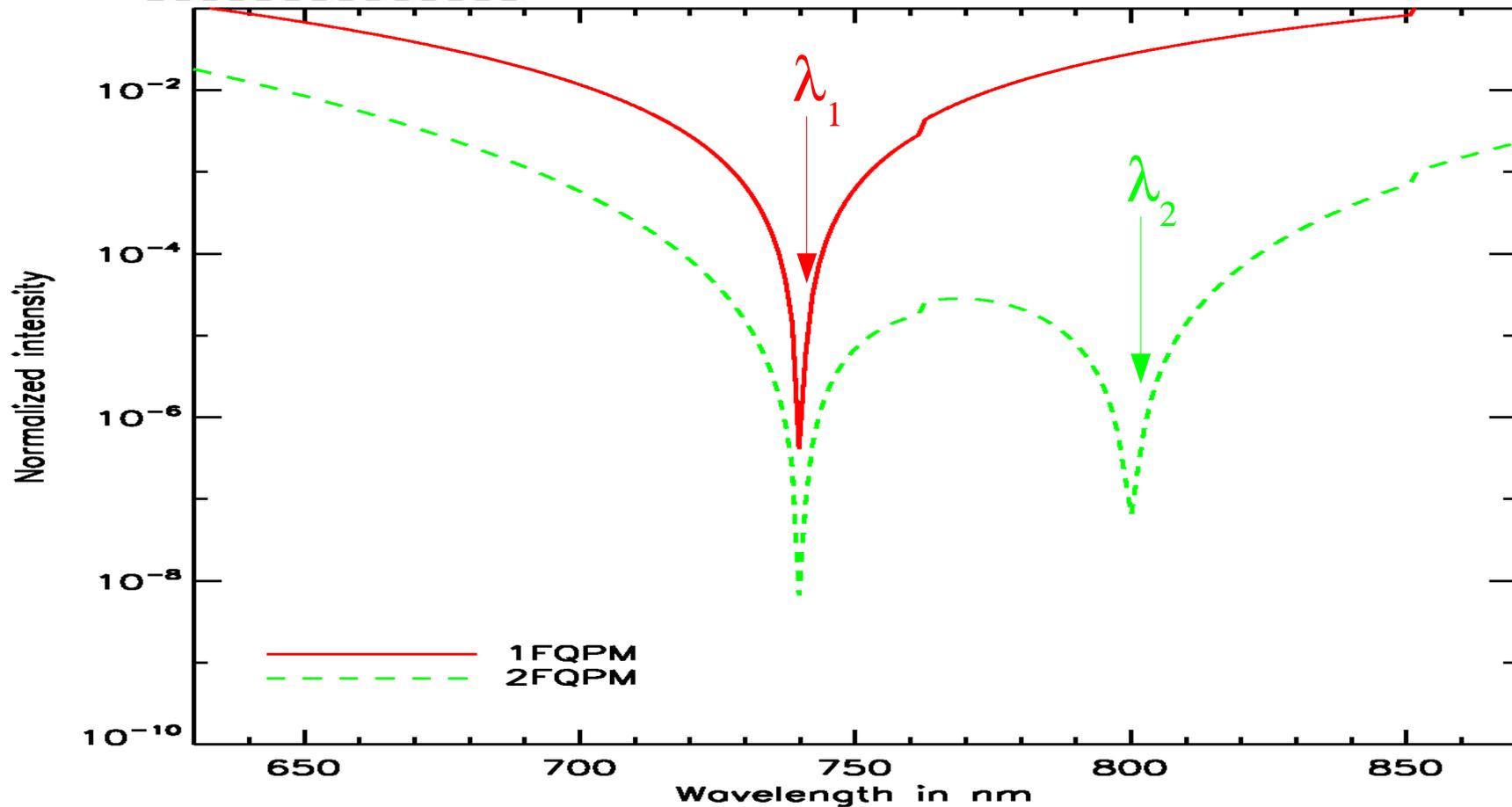
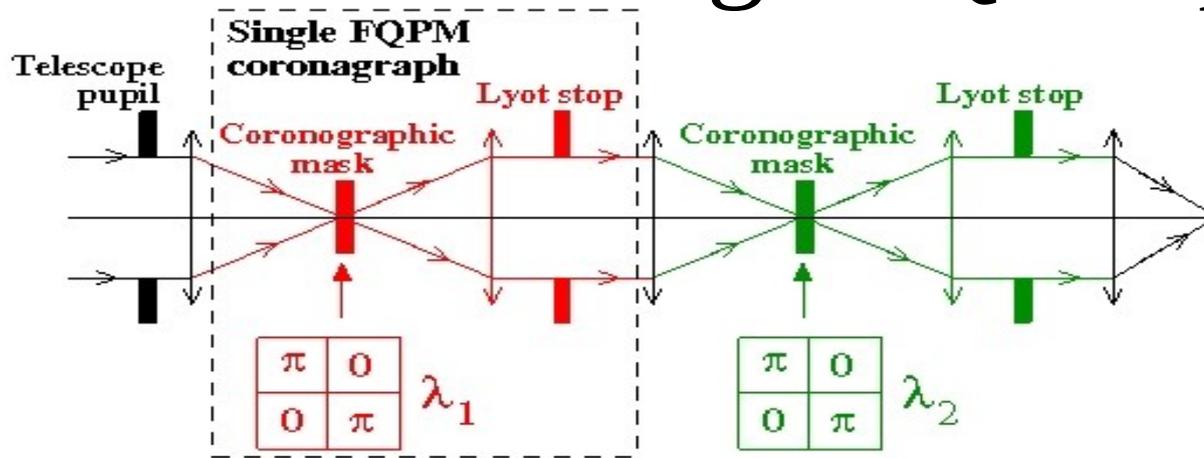
Performance with obscuration

Conclusions

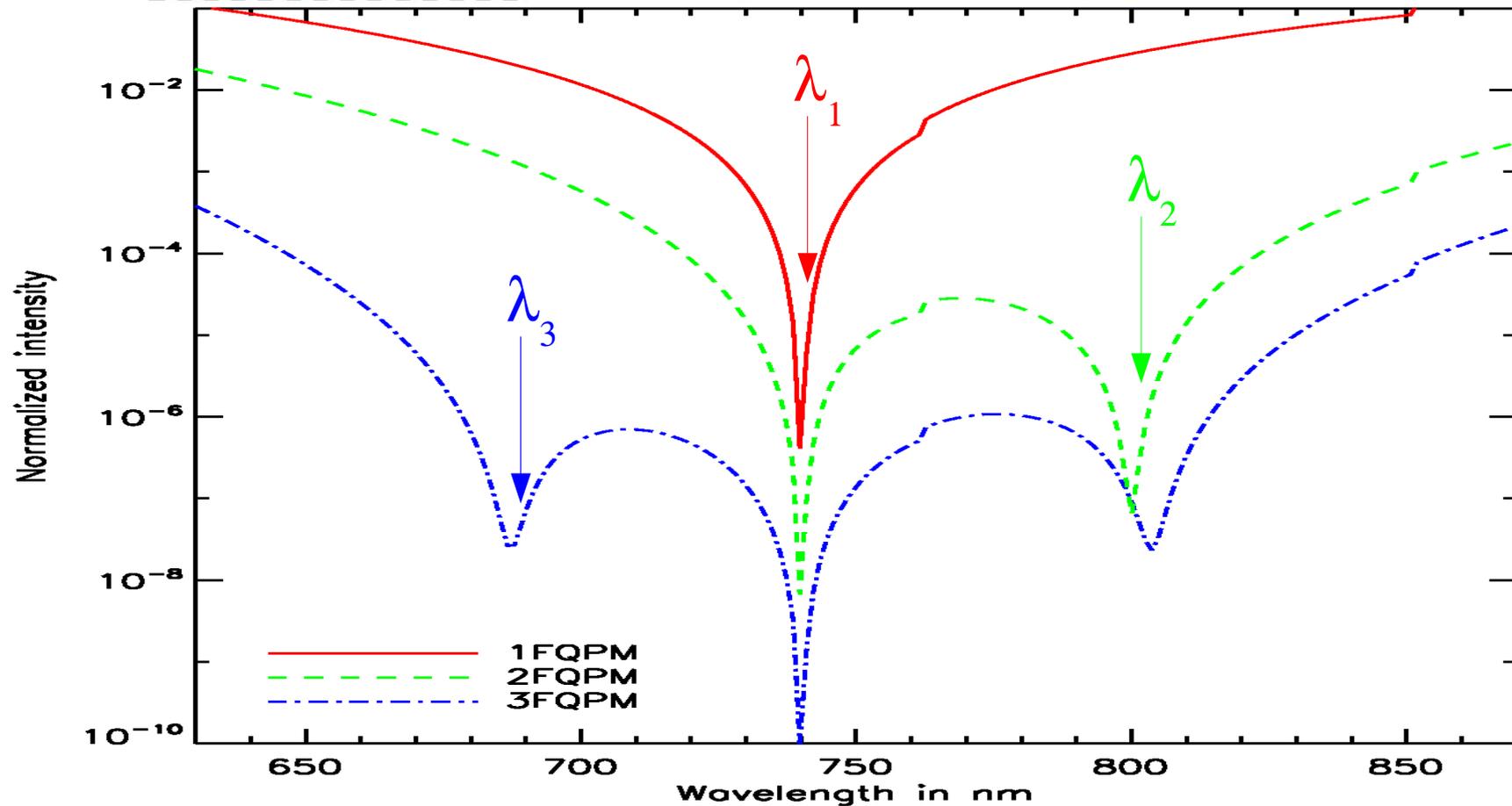
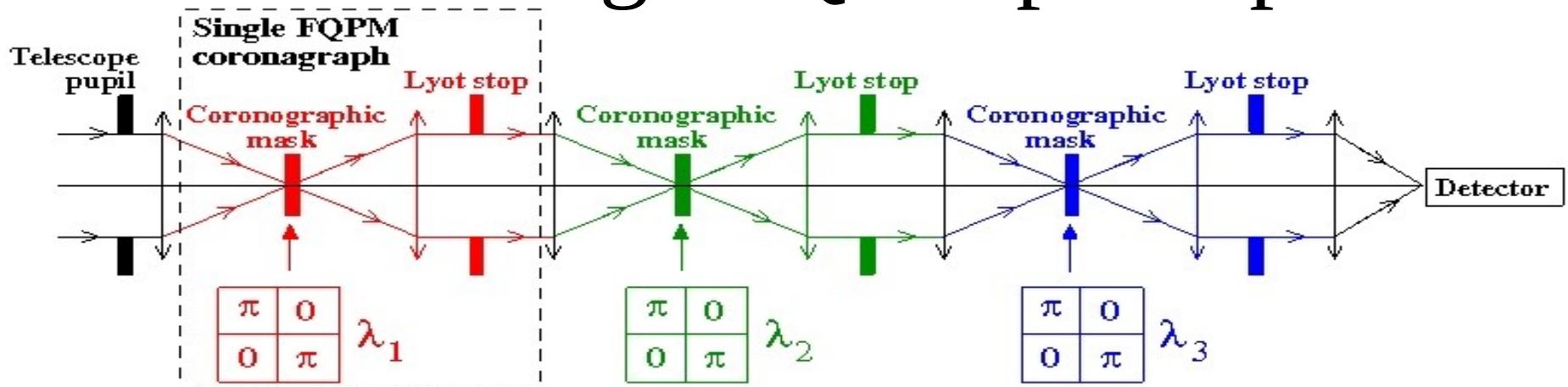
Multi-stage FQPM principle



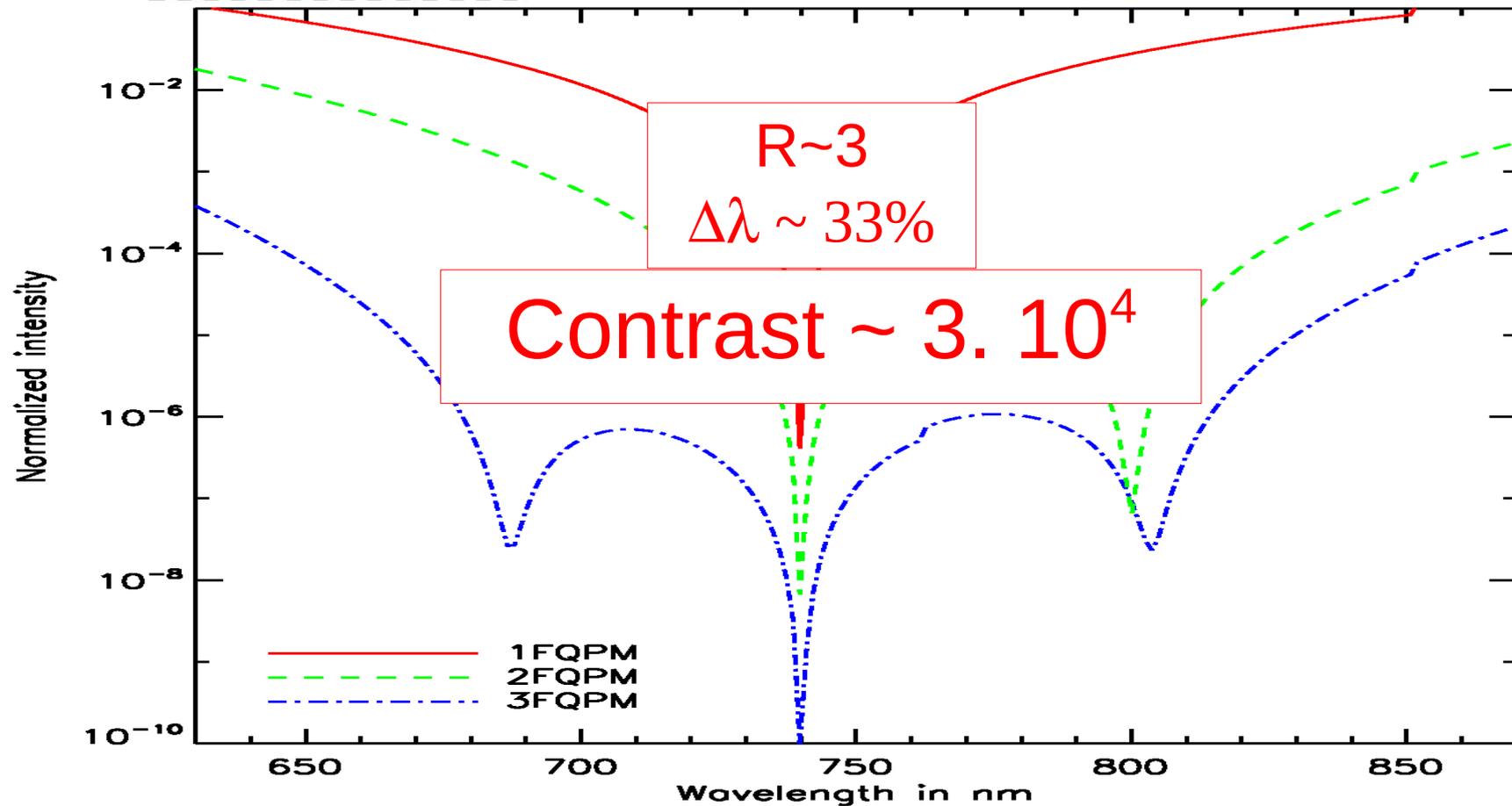
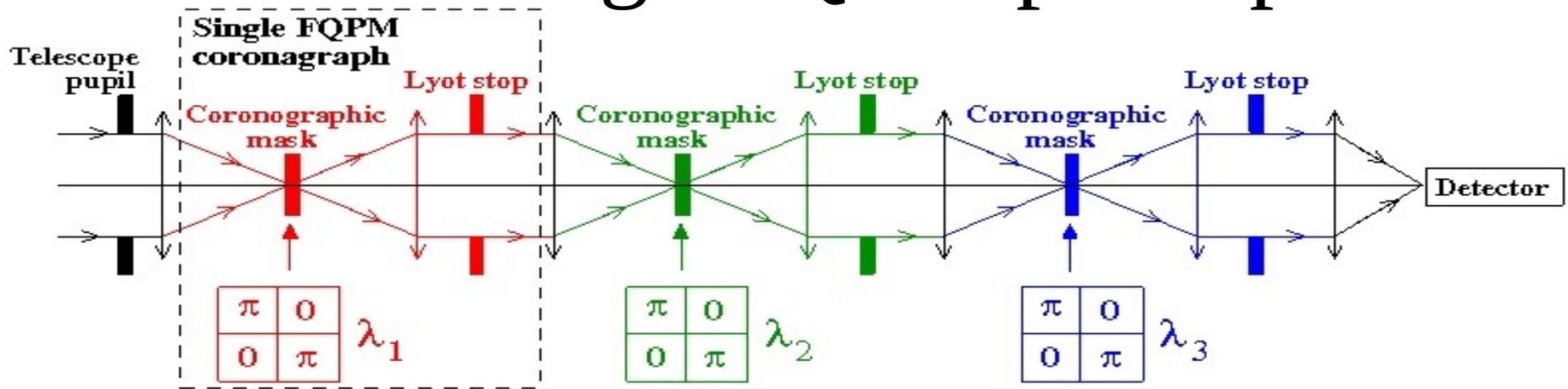
Multi-stage FQPM principle



Multi-stage FQPM principle



Multi-stage FQPM principle



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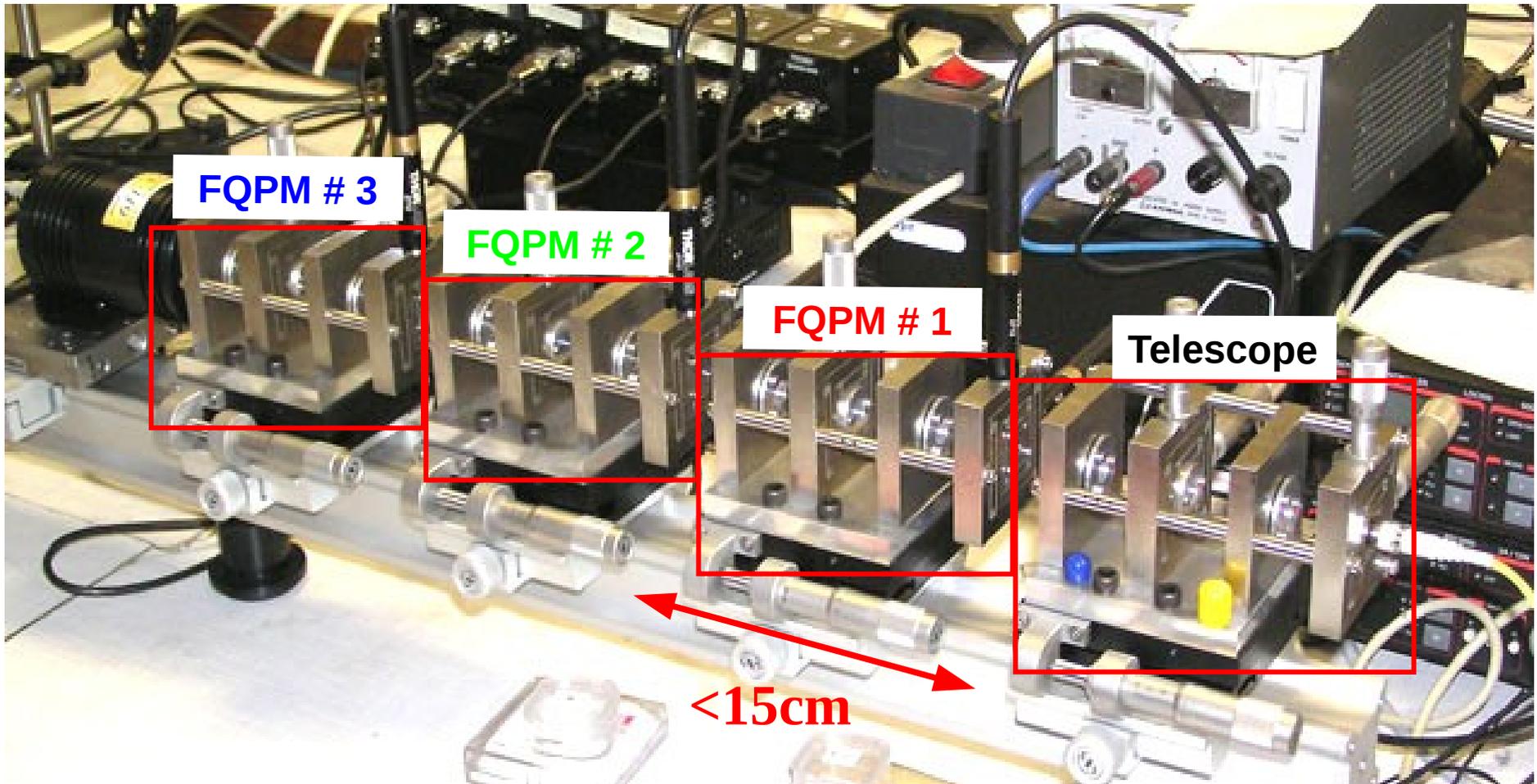
Conclusions

MFQPM prototype

630nm to 870nm ($\Delta\lambda=240\text{nm}$, 32% bandwidth)

Aperture : F/D = 40

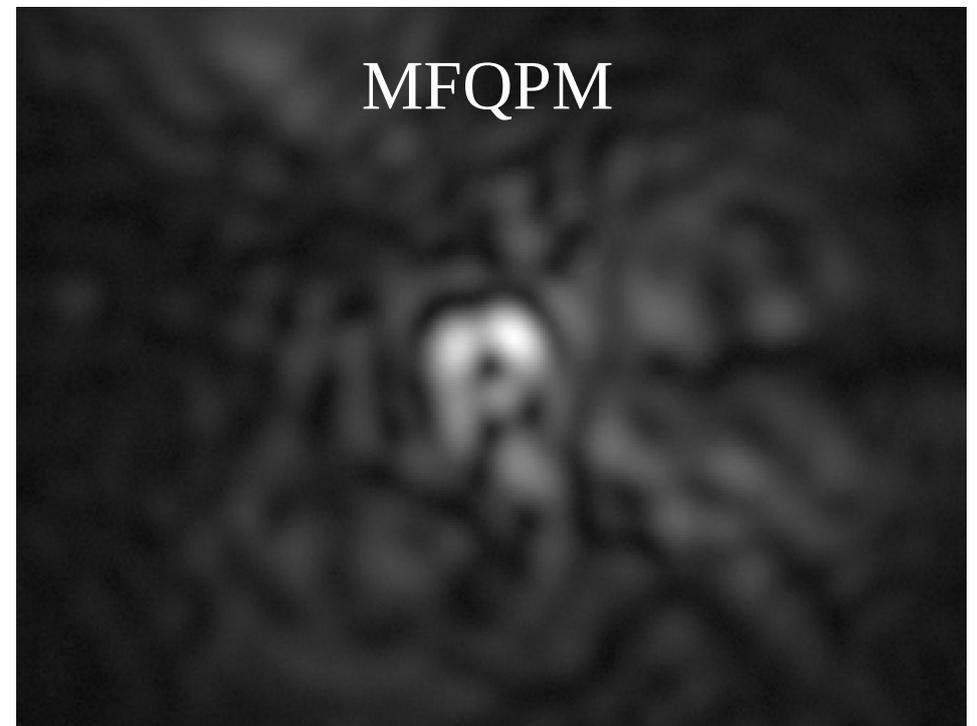
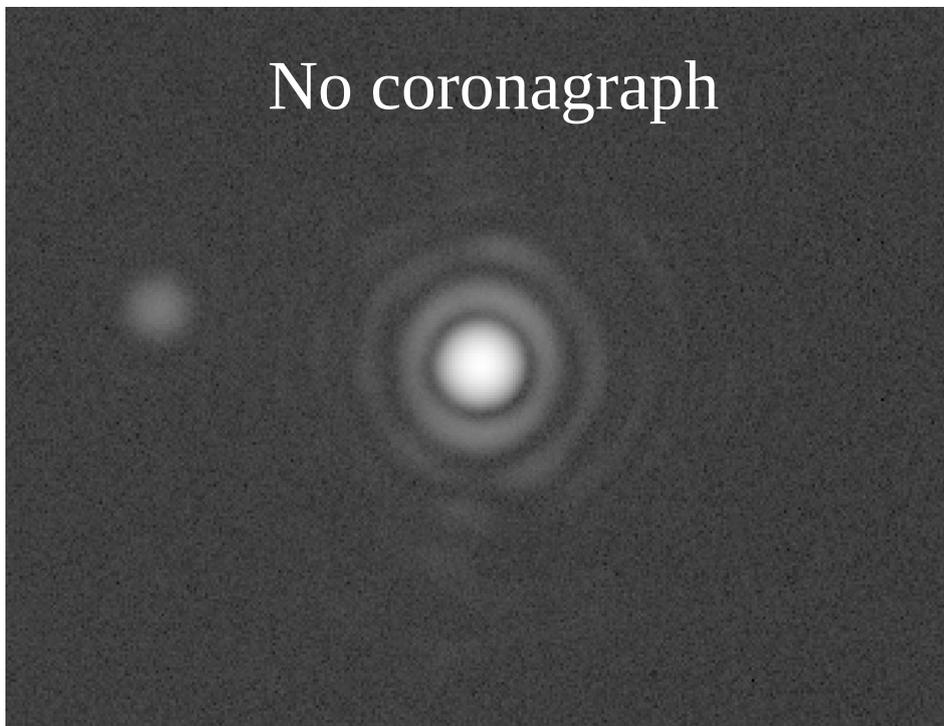
$\lambda_3=800\text{nm}$ $\lambda_2=690\text{nm}$ $\lambda_1=740\text{nm}$



Full pupil R=3 : images

- 630 to 870nm → $\Delta\lambda=32\%$
- Aperture F/D=40
- Throughput = 72%

Total energy rejection ~ 2000
→ limited by speckles (no calibration)

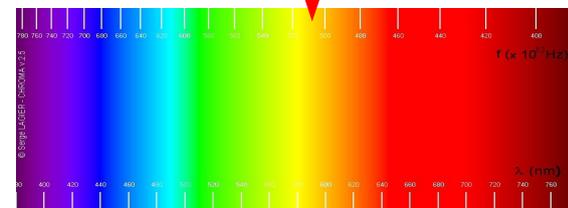
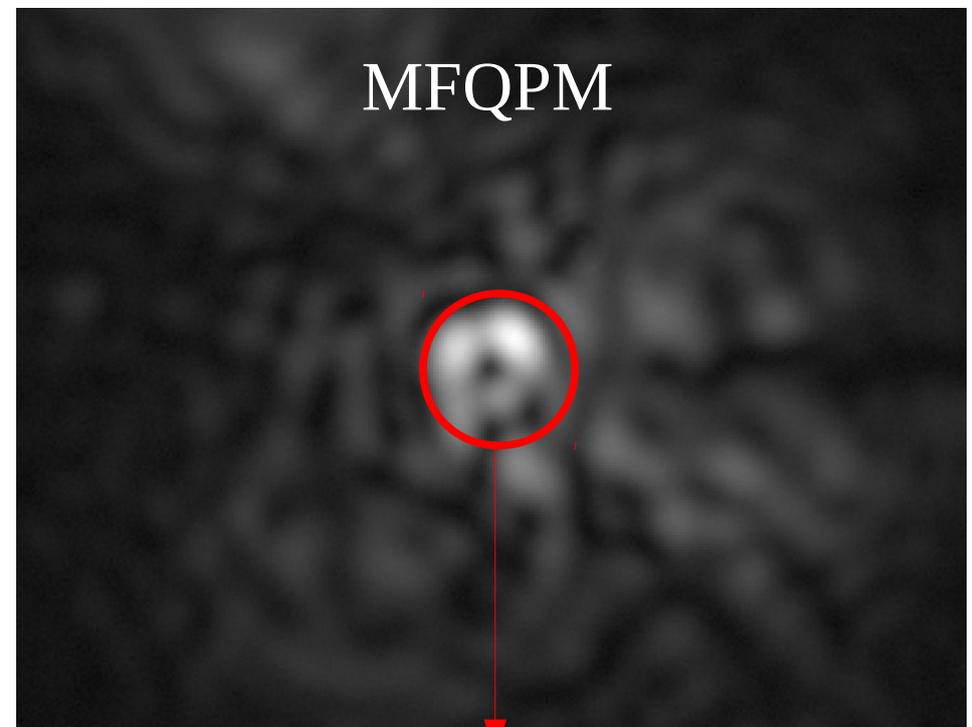
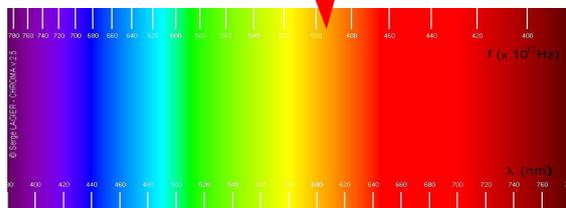
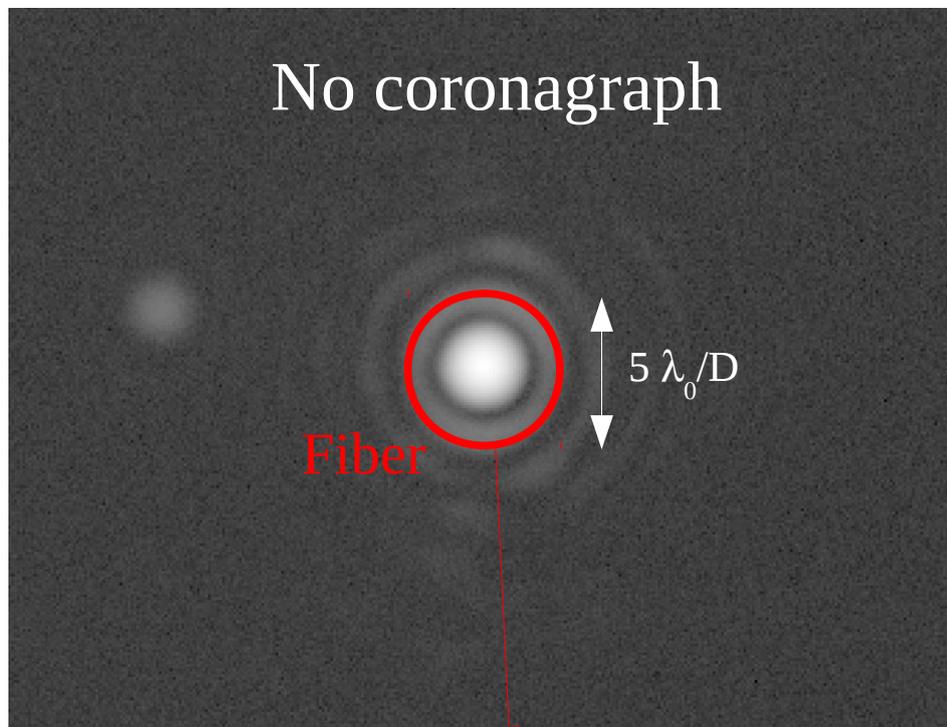


Different scales

Full pupil R=3 : images

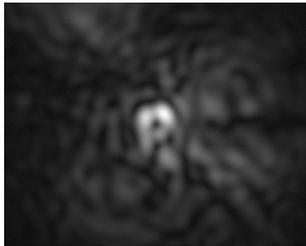
- 630 to 870nm $\rightarrow \Delta\lambda=32\%$
- Aperture F/D=40
- Throughput = 72%

Total energy rejection ~ 2000
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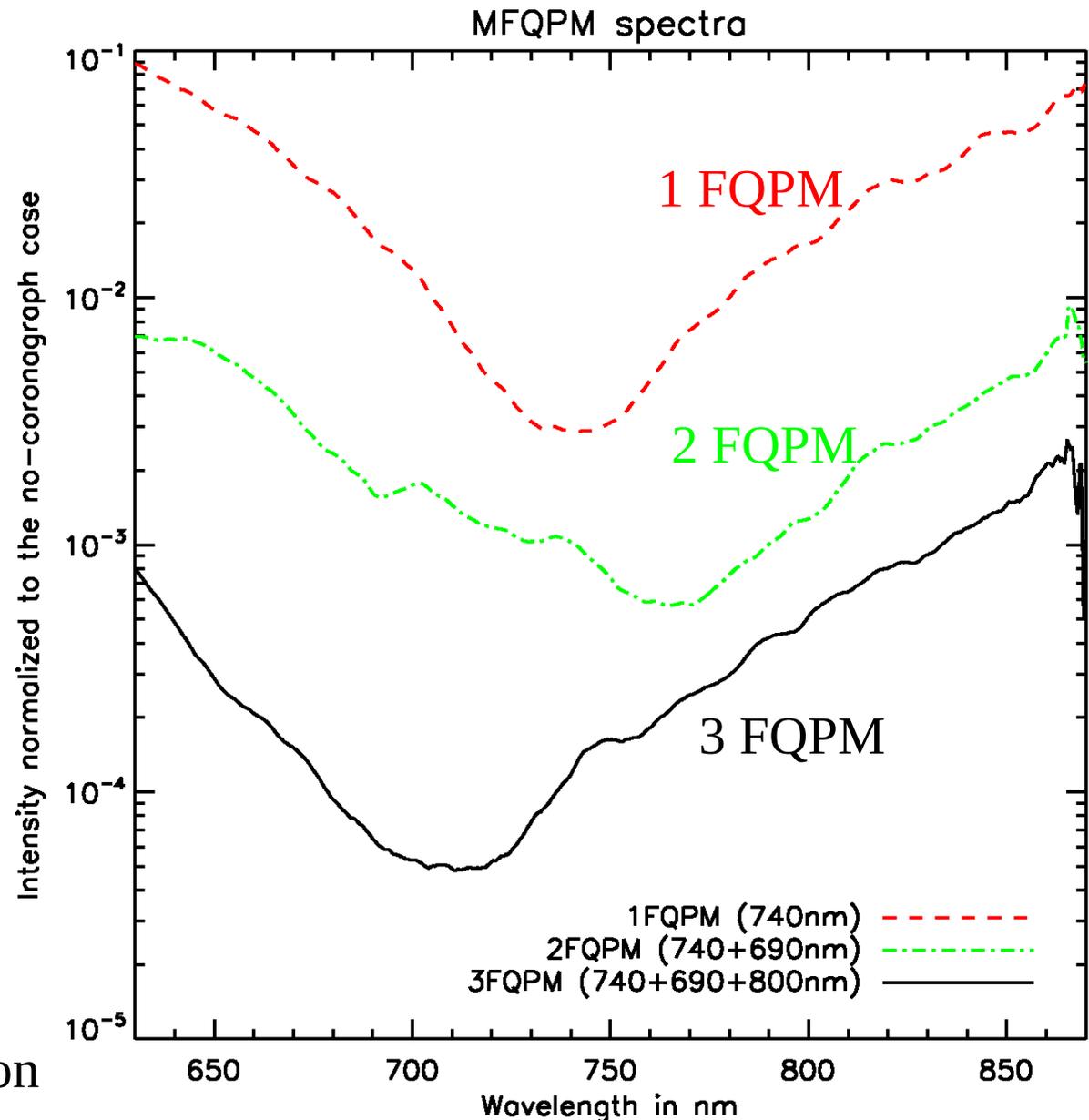
Full pupil R=3 : spectra

- Throughput = 72%
- Aperture : F/D=40
- Fiber of $5\lambda_0/D$ diameter put on-axis



Improvements :

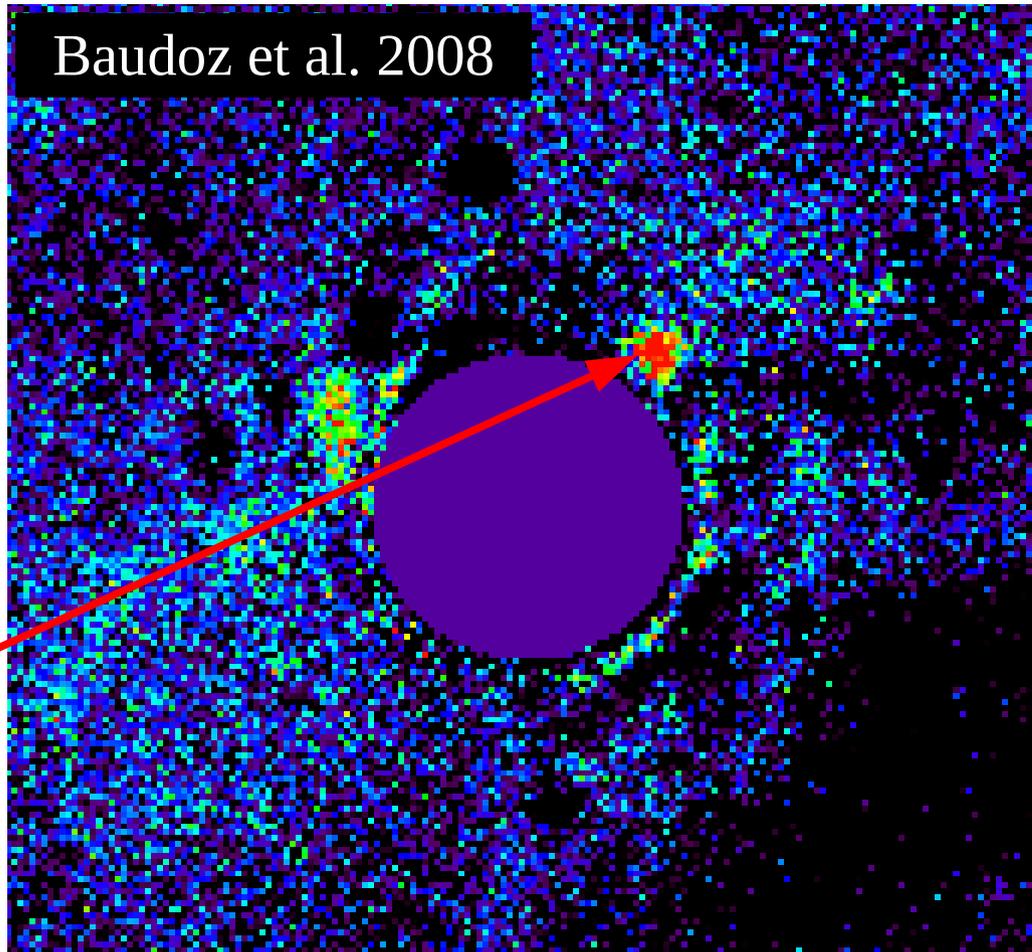
- Flatness of spectra
 - focal masks centering
- Contrast
 - speckle limitation
 - better optics
 - active correction/calibration



Full pupil R=4.8

- Throughput = 56%
- Aperture : F/D=250-300
- Differential imaging
(factor 10 on raw contrast)

Contrast $1.5 \cdot 10^8$
at $4.5\lambda_0/D$
SNR=4.5



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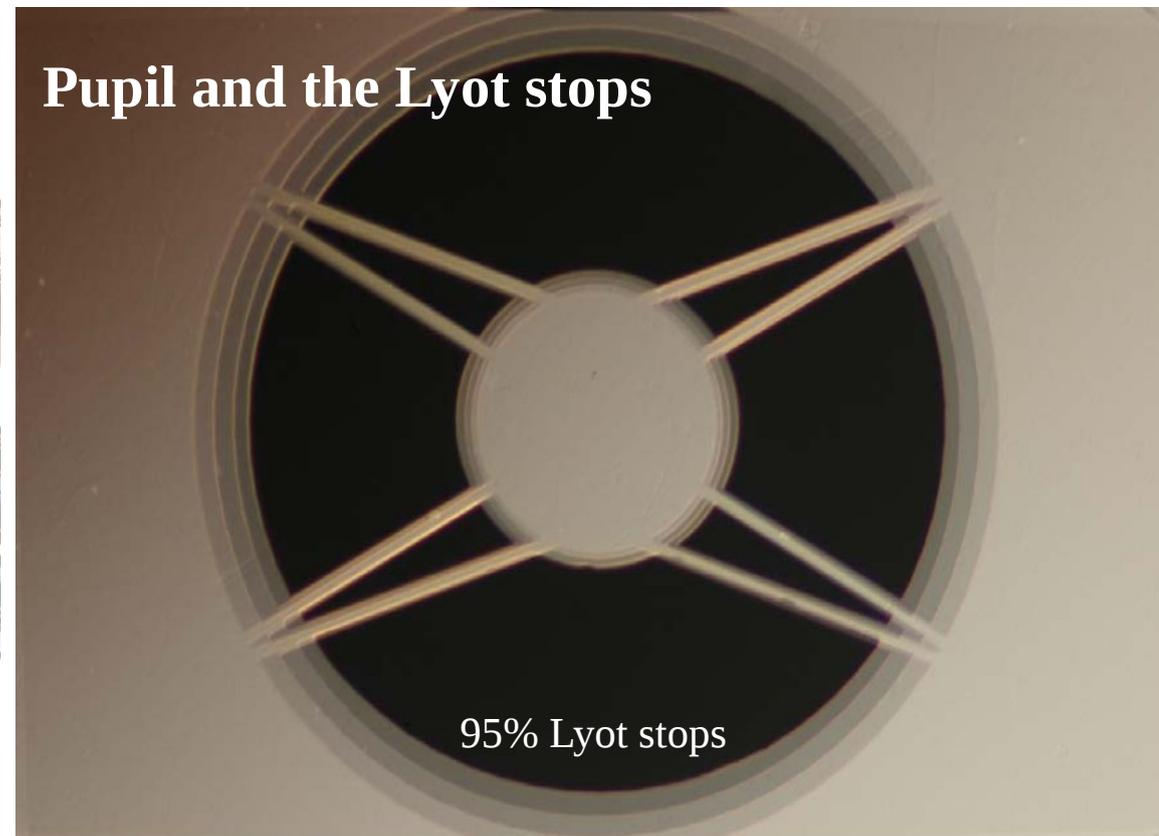
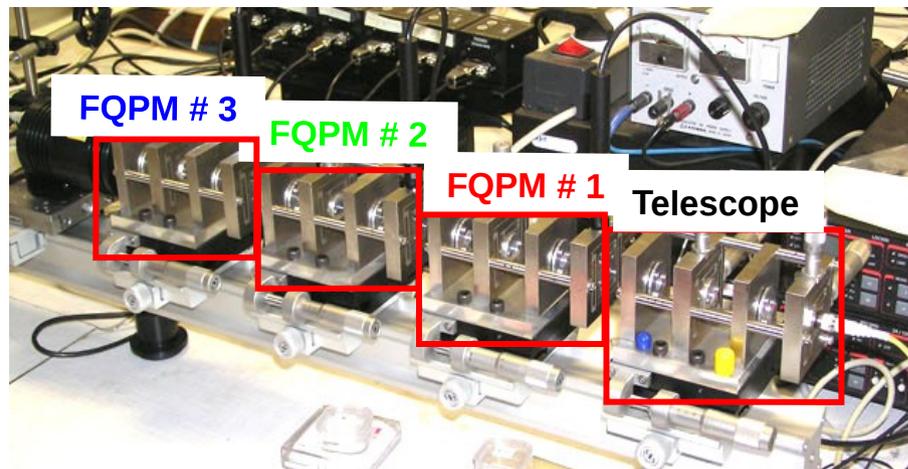
Performance in full pupil

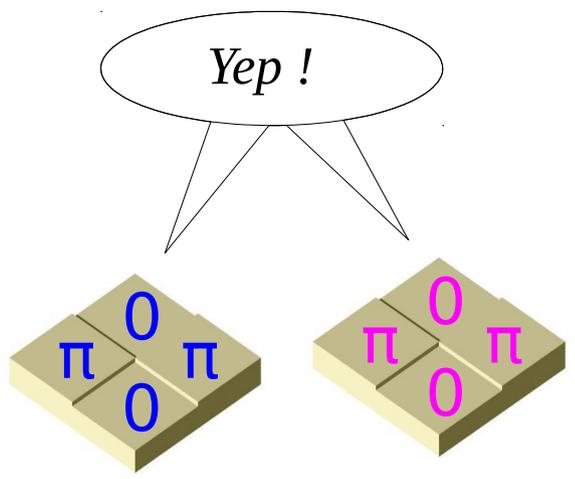
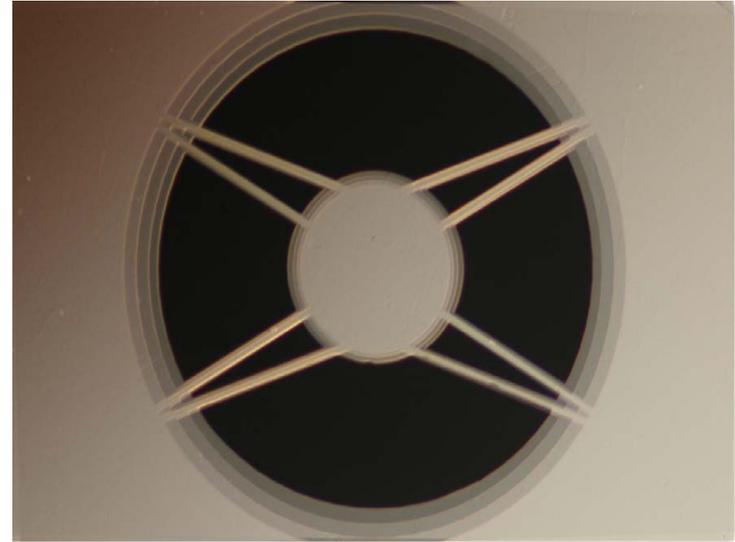
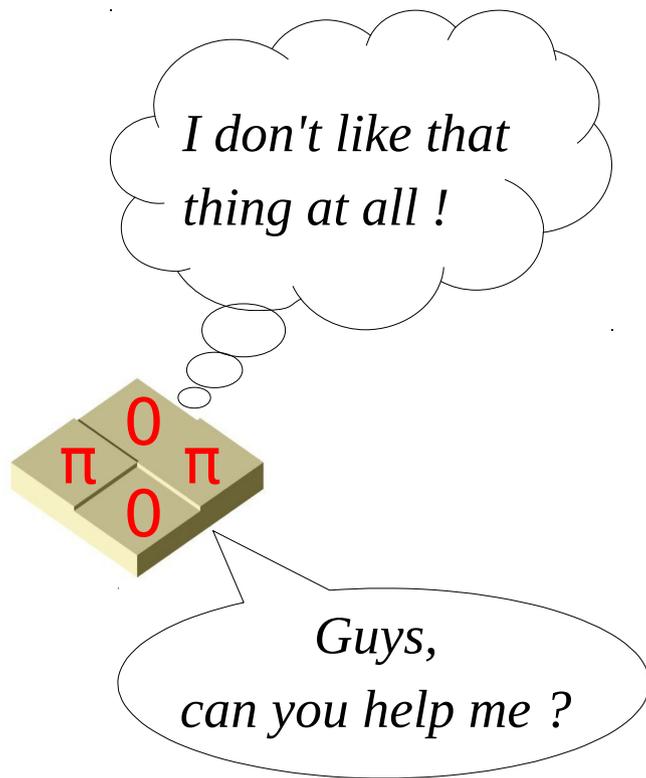
Performance with obscuration

Conclusions

Obstructed pupil R=3

- 630 to 870nm $\rightarrow \Delta\lambda=32\%$
- Aperture : F/D=40
- E-ELT pupil
- Throughput = 86%



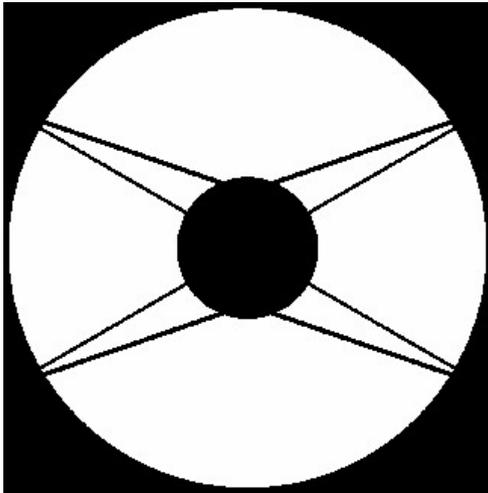


Obstructed pupil R=3 : pupil images

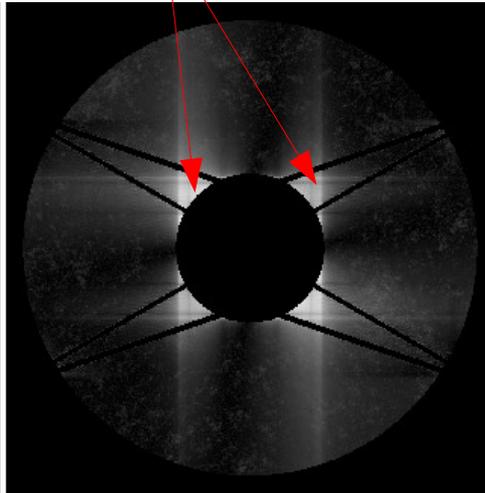
Central obstruction effects

Leaks are stopped

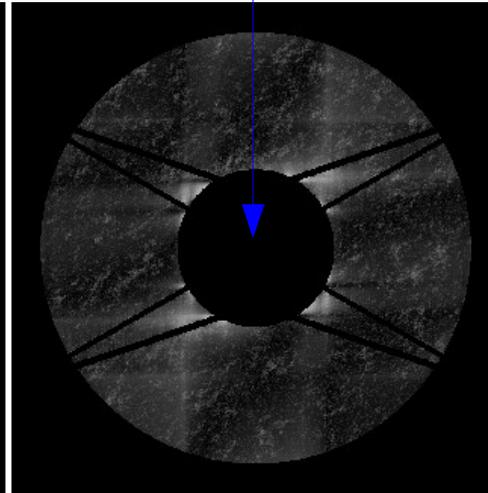
Throughput = 86%



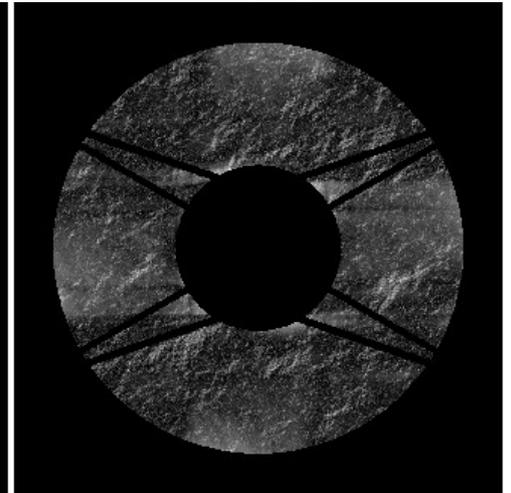
No coronagraph



1st Lyot stop



2nd Lyot stop



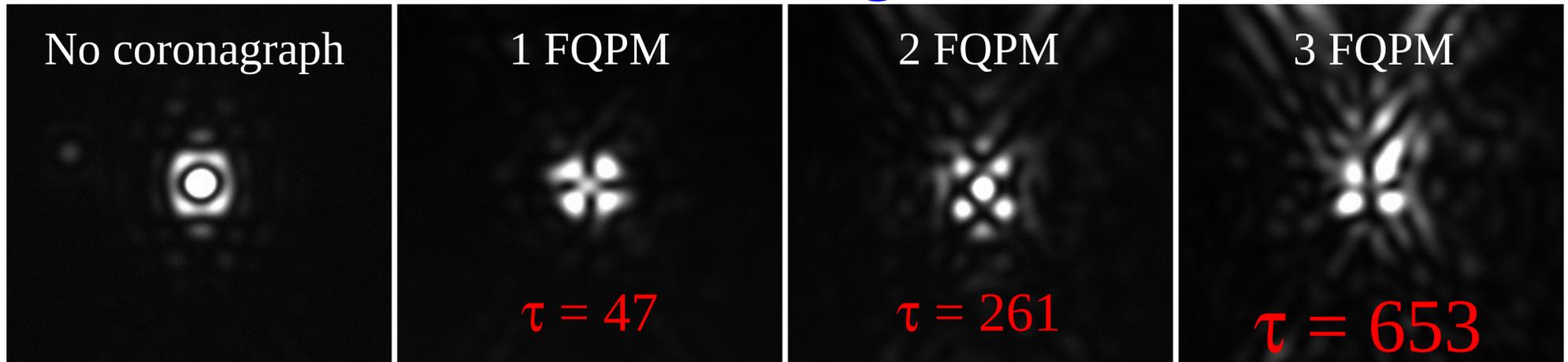
3rd Lyot stop

Obstructed pupil R=3 : images

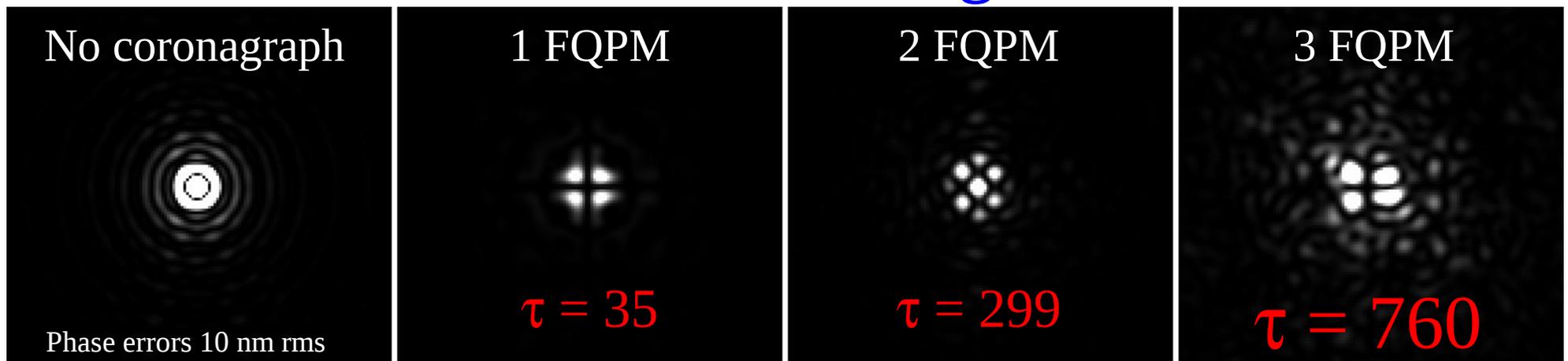
630 to 870nm $\rightarrow \Delta\lambda=32\%$ **Throughput = 86%** No speckle calibration

τ = total rejection

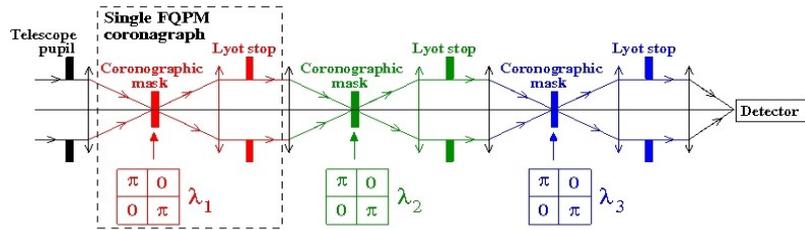
Lab images



Numerical images



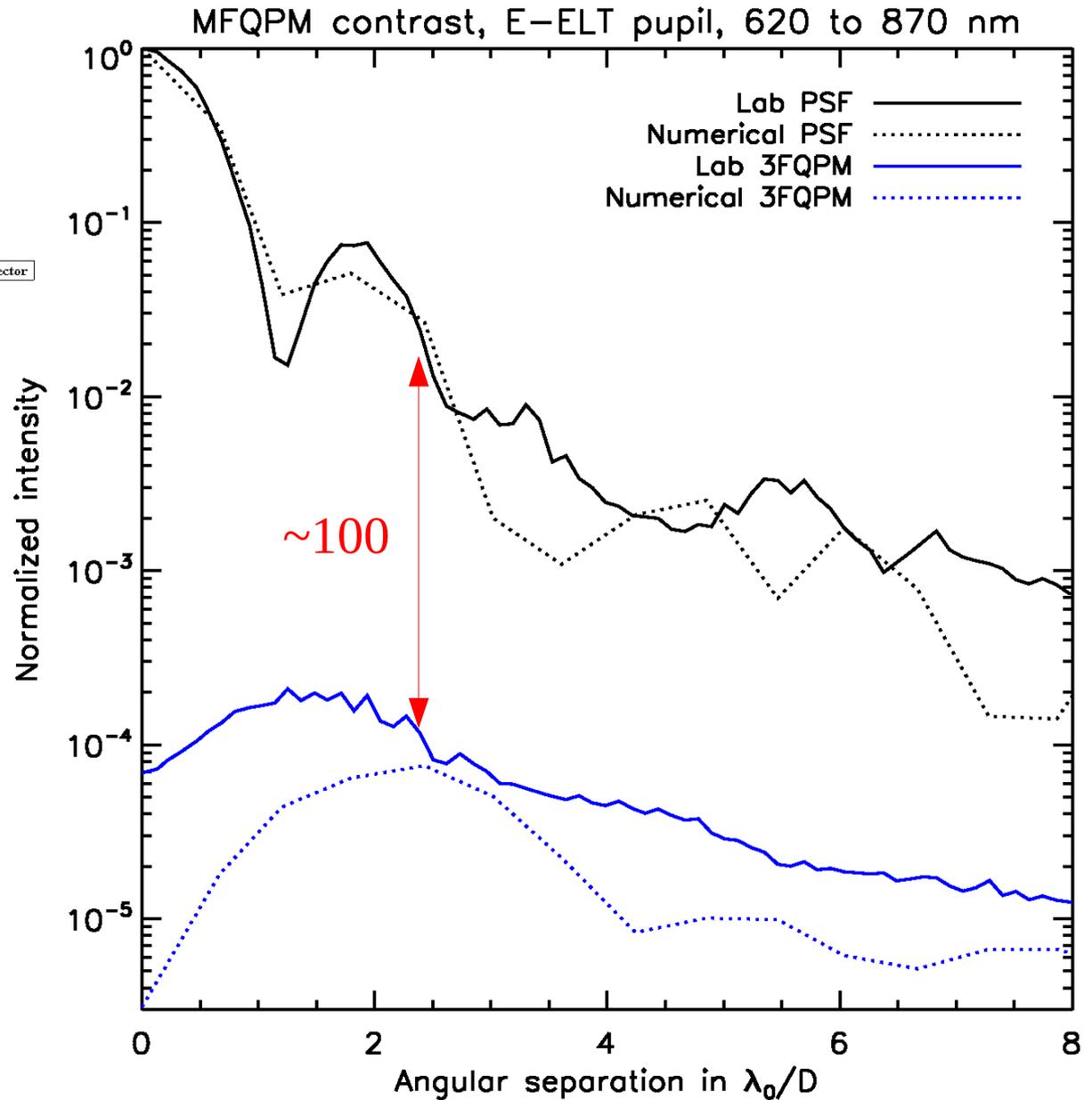
Obstructed pupil R=3 : images



630 to 870nm $\rightarrow \Delta\lambda=32\%$

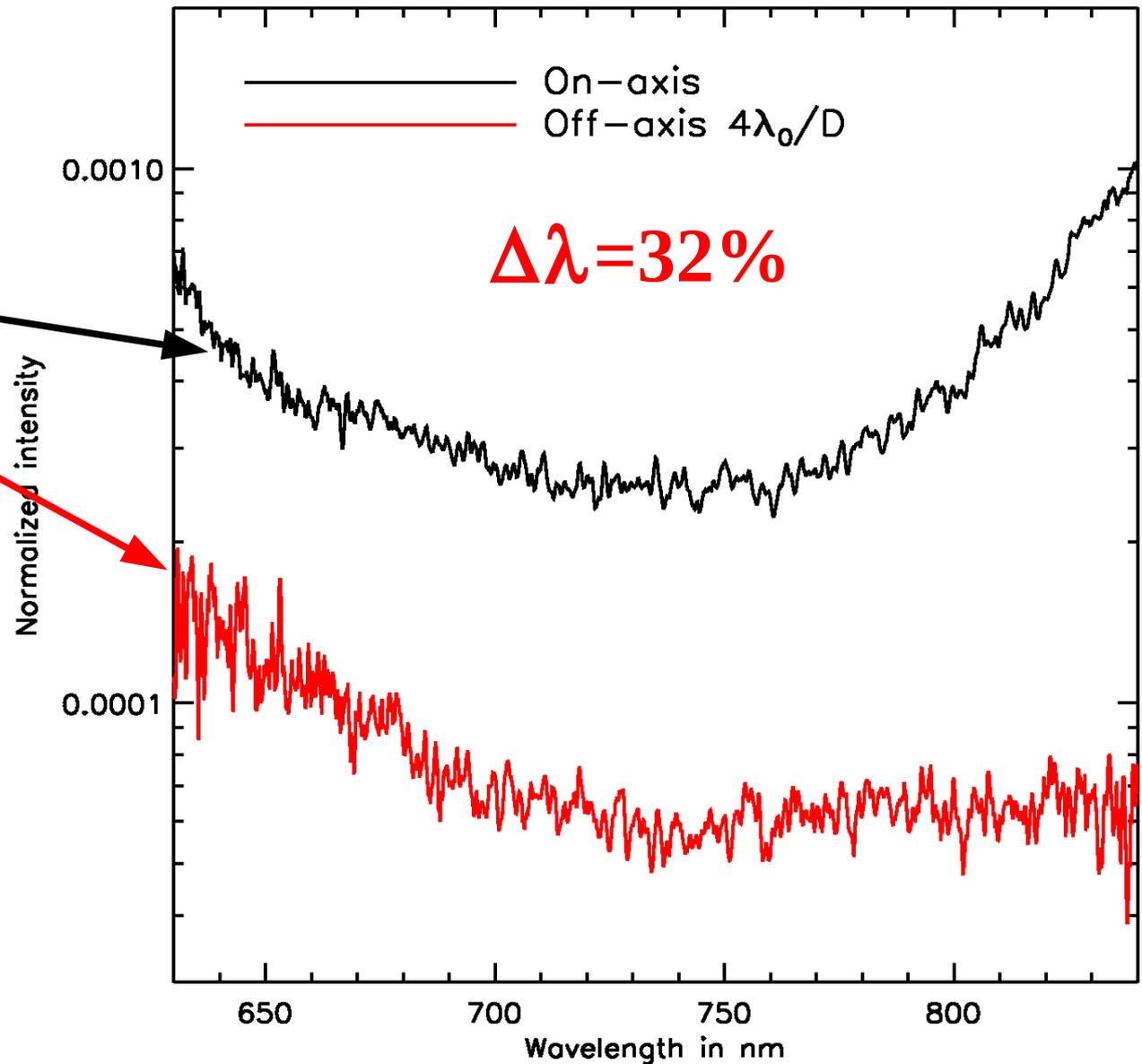
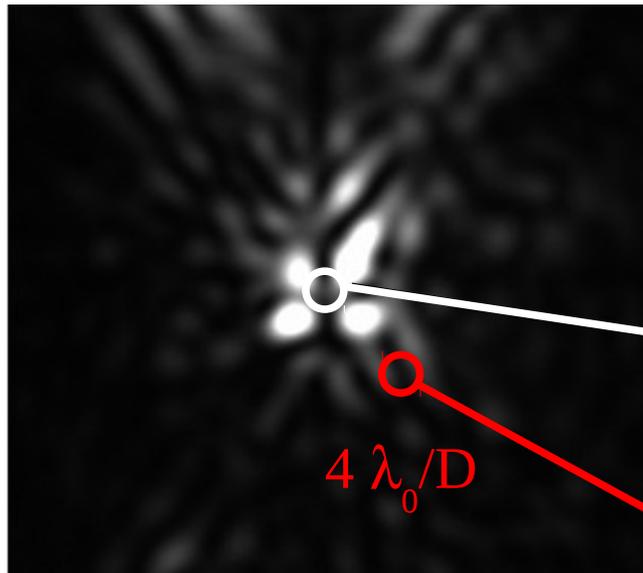
Throughput = 86%

No speckle calibration



Obstructed pupil R=3 : spectra

MFQPM E-ELT T=86% $\Delta\lambda=32\%$



Fiber diameter $\sim 1\lambda_0/D$

No speckle calibration

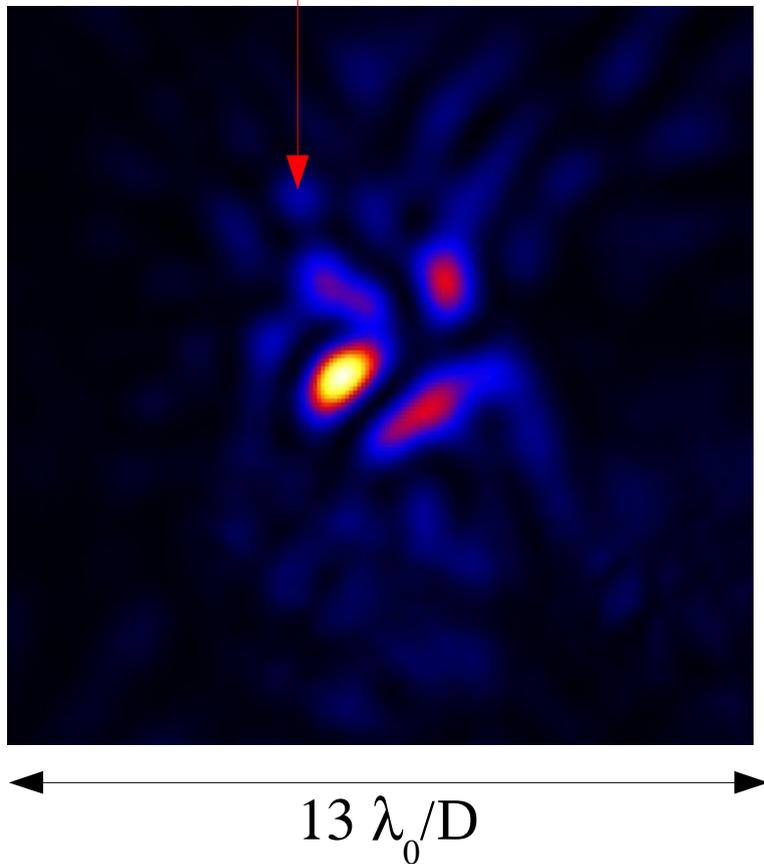
Obstructed pupil R=3 : lab planet

630 to 870nm $\rightarrow \Delta\lambda=32\%$

E-ELT pupil

Throughput = 86%

Lab planet : $1.5 \cdot 10^{-4}$ at $3.47 \lambda_0/D$



Detected with SNR=2

No speckle calibration

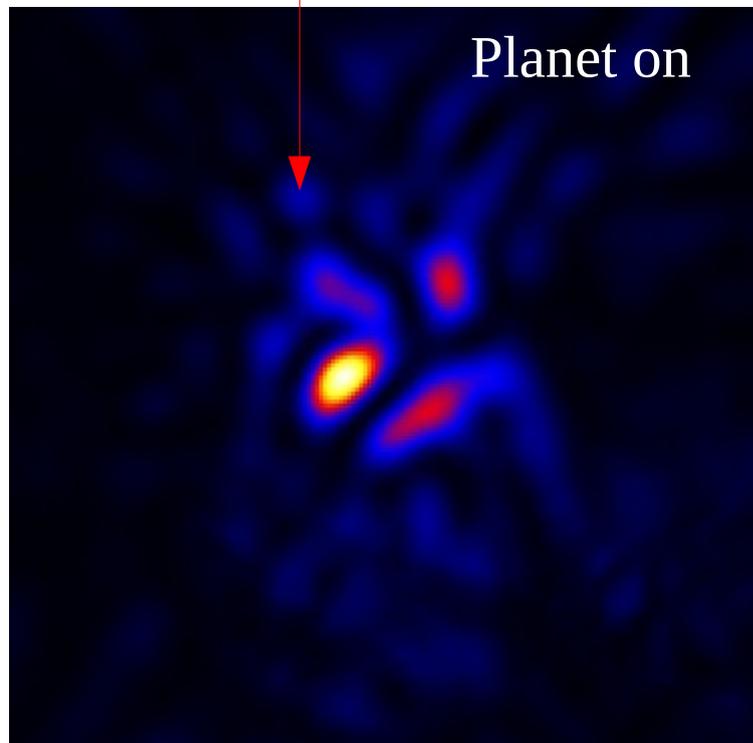
Obstructed pupil R=3 : lab planet

630 to 870nm $\rightarrow \Delta\lambda=32\%$

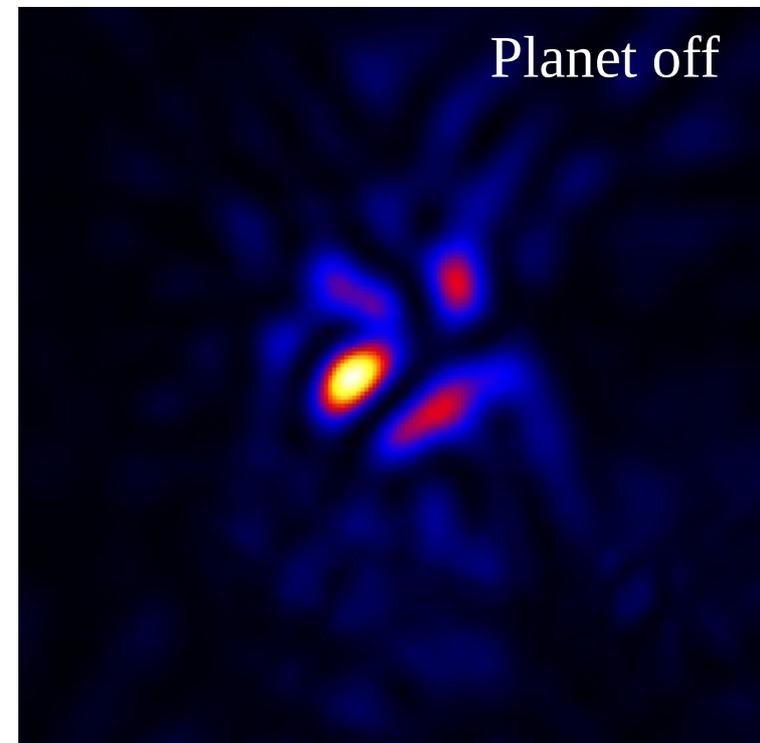
E-ELT pupil

Throughput = 86%

Lab planet : $1.5 \cdot 10^{-4}$ at $3.47 \lambda_0/D$



$13 \lambda_0/D$



$13 \lambda_0/D$

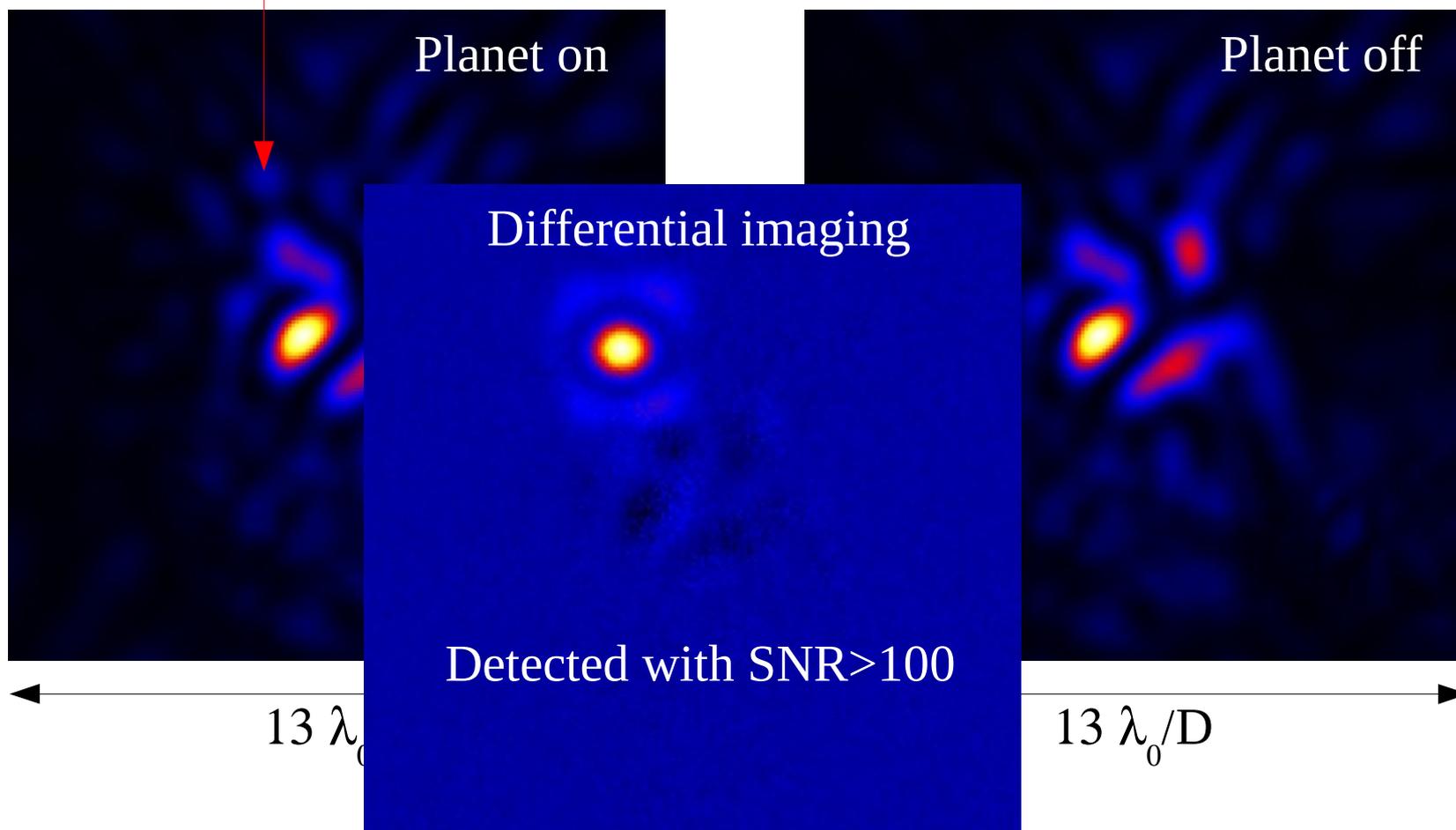
Obstructed pupil R=3 : lab planet

630 to 870nm $\rightarrow \Delta\lambda=32\%$

E-ELT pupil

Throughput = 86%

Lab planet : $1.5 \cdot 10^{-4}$ at $3.47 \lambda_0/D$



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Conclusions

- MFQPM **very easy to build** (works at every λ) already specified for space (JWST)
- Full pupil, F/40, 32% bandwidth, T=72% **2000 total energy rejection**
Full pupil, F/250, 20% bandwidth, T=56% **10^8 contrast at $4.5 \lambda_0/D$**
- **Obstructed pupil (E-ELT), F/40, 32% bandwidth, T=86%**
650 total energy rejection
 $1.5 \cdot 10^{-4}$ lab planet at $3.5 \lambda_0/D$ detection

Ongoing works :

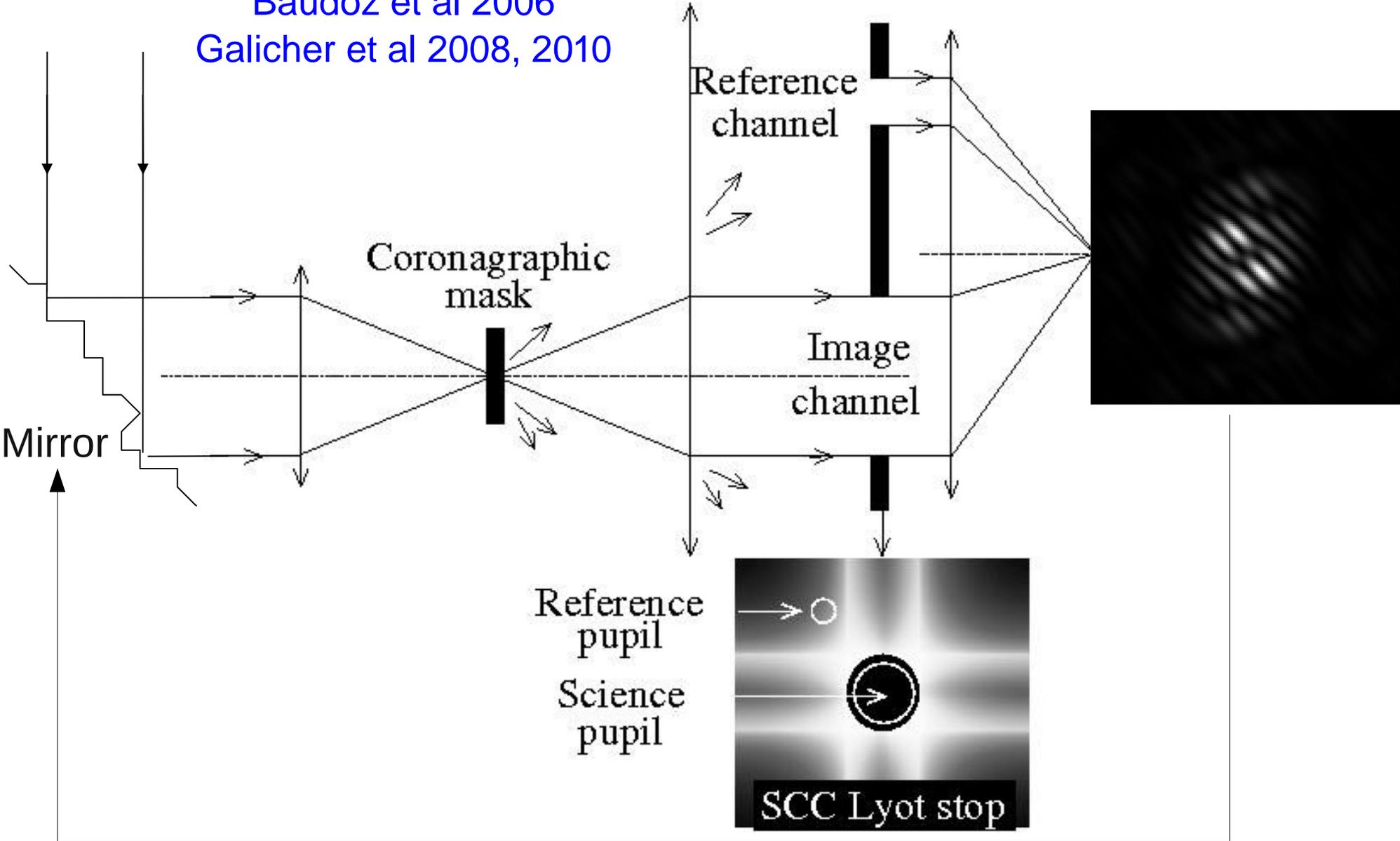
- **New prototype** with less optic aberrations → flatten spectra
- **Association with a speckle calibrator/killer (Self-coherent camera)**

See Marion Mas talk tomorrow

Thanks

Self-coherent camera + four quadrant phase mask coronagraph

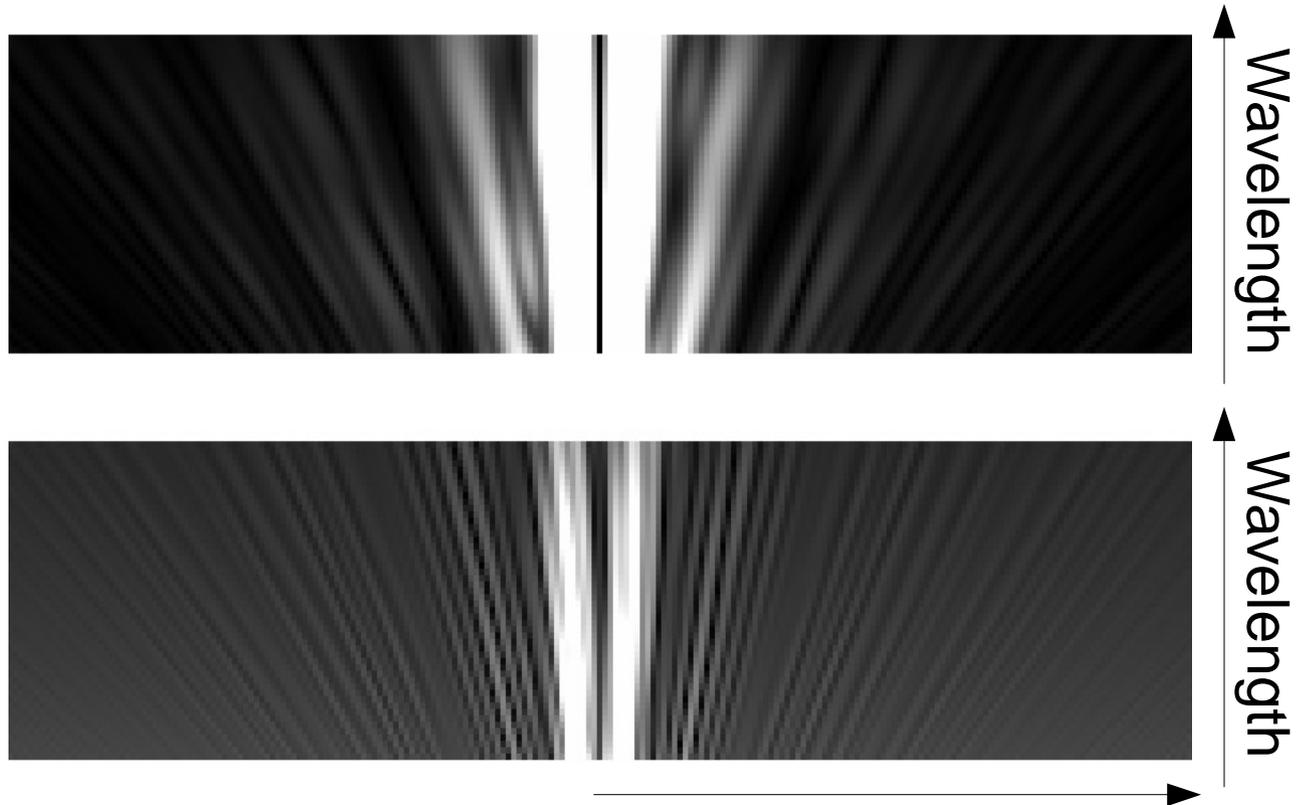
Baudoz et al 2006
Galicher et al 2008, 2010



Integral Field Spectrometer + Self-coherent camera

Aberrations = function of λ (Fresnel propagation)

**Spectral
deconvolution
strongly limited**



**One solution :
SCC-IFS**

MFQPM prototype

